

Theoretical paper

Notebook: myao2018

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Location: 5°21'30 N 100°18'18 E

Theoretical

1) We use Stefan-Boltzmann law

$$E = \sigma T^4 \text{ J} \cdot \text{m}^{-2} \cdot \text{s}^{-1} \quad (40)$$

where E : the total energy radiated per unit ~~area~~ surface area of a black body across all wavelengths per unit time

$$\sigma: \text{Stefan-Boltzmann constant} \\ = 5.670373 \times 10^{-8} \text{ W} \cdot \text{m}^{-2} \cdot \text{K}^{-4}$$

For the photosphere, we have

$$E_1 = \sigma T_1^4 = \sigma (5,800 \text{ K})^4 \quad (20)$$

For the sunspot, we have

$$E_2 = \sigma T_2^4 = \sigma (4,200 \text{ K})^4 \quad (20)$$

∴ each square meter of the photosphere is brighter than each square meter of the sunspot by

$$\frac{E_1}{E_2} = \frac{\sigma T_1^4}{\sigma T_2^4} = \left(\frac{5,800 \text{ K}}{4,200 \text{ K}} \right)^4 = 3.636 \text{ times} \quad (20)$$

(2) We use Pogson's Equation

$$M = m - 5 \log_{10} d + 5 \quad (40)$$

where M : absolute magnitude of star

m : apparent magnitude of star

d : distance of star from Sun in parsecs

(a) We have, for the first star,

$$M = m_1 - 5 \log_{10} d_1 + 5 \quad \text{--- (1)}$$

and for the second star,

$$M = m_2 - 5 \log_{10} d_2 + 5 \quad \text{--- (2) (20)}$$

$$\text{and } \frac{d_1}{d_2} = 1,500$$

$$\begin{aligned} \therefore m_1 - 5 \log_{10} d_1 + 5 &= m_2 - 5 \log_{10} d_2 + 5 \\ \therefore m_1 - m_2 &= \log_{10} \left(\frac{d_1}{d_2} \right)^5 = \log_{10} (1,500)^5 \\ &= 15.88 \quad (20) \end{aligned}$$

(b) The larger apparent magnitude is $m_1 = 15.88 + m_2$ for the star that is further away
(20)

(3)

(a) From Kepler's Third Law

$$T^2 \propto r^3$$

$$= \frac{4\pi^2 r^3}{GM} \quad (40)$$

where T : period around the Sun r : semi-major axis of space probe
in its orbit around the Sun G : universal gravitational constant M : mass of Sun

We also have, $r = \frac{r_p + r_a}{2}$ (20)

where r_p : perihelion distance r_a : aphelion distance

$$r = \frac{0.4 \text{ AU} + 5.4 \text{ AU}}{2} = 2.9 \text{ AU}$$

$$= 2.9 \times 1.496 \times 10^8 \text{ km}$$

$$= 4.338 \times 10^{11} \text{ m}$$

$$T^2 = \frac{4\pi^2 (4.338 \times 10^{11} \text{ m})^3}{(6.674 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2})(1.989 \times 10^{30} \text{ kg})}$$

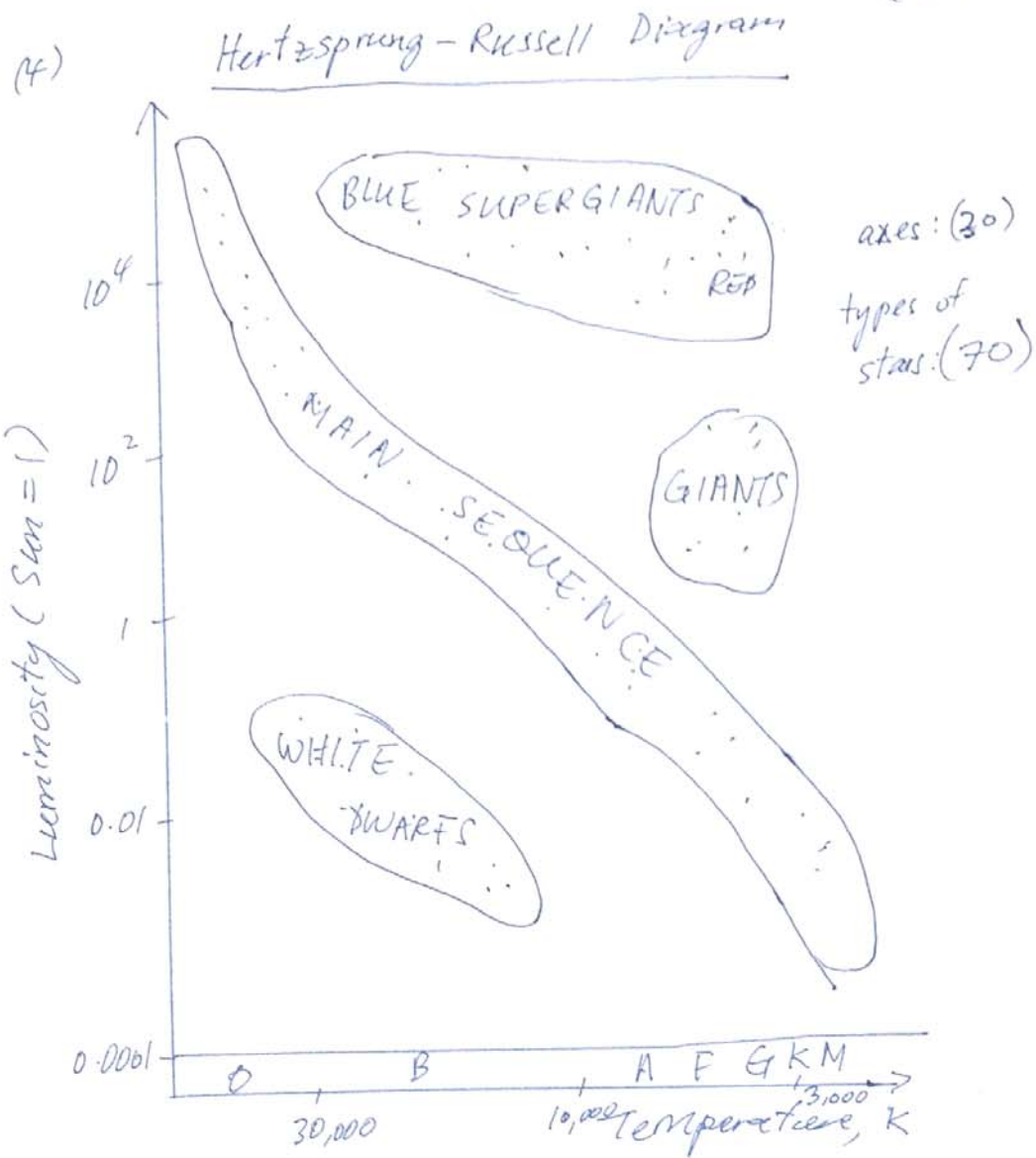
$$= 2.428 \times 10^{16} \text{ s}^2$$

and $T = \sqrt{2.428 \times 10^{16} \text{ s}^2} = 155,835,718 \text{ s}$

$$= \frac{155,835,718}{3600 \times 24 \times 365.25}$$

$$= 4.938 \text{ years} \quad (20)$$

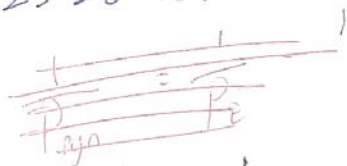
(b) Since the space probe is in orbit around the Sun, its orbit cannot be a parabola or hyperbola. Since its orbit has a perihelion and an aphelion, its orbit cannot be a circle and is therefore an ellipse. (20)



(51) The planet Mars was at opposition with Earth at UTC 13 hours 13 minutes on 27th July 2018. Each Mars opposition is separated from the next opposition by ~~778~~ days, or 2 years, 1 month and 18 days. Therefore the next Mars opposition will occur at

~~13 October 2020~~

23 20 UTC on 13 October 2020 (30)



~~100~~
(50)

$$\frac{1}{P_{syn}} = \frac{1}{P_e} - \frac{1}{P_m} \quad (40)$$

$$= \frac{1}{365.26} - \frac{1}{686.97}$$

$$= 779.965 \text{ Earth days} \quad (30)$$

The Schwarzschild Radius, r_s , is given by

$$r_s = \frac{2GM}{c^2} \quad (40)$$

where G : universal gravitational constant

M : mass of supermassive black hole
 $= 4 \times 10^6 M_\odot$

where $M_\odot$ is the mass of the Sun

c : speed of light

$$\begin{aligned} r_s &= \frac{2(6.674 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2})(4 \times 10^6 \times 1.989 \times 10^{30} \text{ kg})}{(2.9979 \times 10^8 \text{ m s}^{-1})^2} \\ &= 1.182 \times 10^{10} \text{ m} \\ &= 11.82 \text{ billion km} \end{aligned} \quad (60)$$

7) (a) We use Pogson's Equation

$$M = m - 5 \log_{10} d + 5 \quad (20)$$

where M : absolute magnitude of star

m : apparent magnitude of star

d : distance of star from Sun in parsecs

Putting in the values, we have

$$2.1 = 7.5 - 5 \log_{10} d + 5$$

Rearranging,

$$d = 10^{2.08} \text{ parsecs}$$

$$= 120.226 \text{ parsecs}$$

$$= 120.226 \times 3.26 \text{ light-years}$$

$$= 391.938 \text{ light years} \quad (30)$$

(b) The luminosity, L , of a star is the total amount of energy emitted per unit time by a star, measured in watts. (20)

$$L = \sigma A T^4$$

where σ : Stefan-Boltzmann constant

A : total surface area of star

T : absolute temperature of surface of star

$$= (A)(\sigma T^4) = (A)(5.670 \times 10^{-8})(9,500\text{K})^4$$

$$= (A \text{ m}^2)(4.618 \times 10^8 \text{ watts})$$

$$= (A)(4.618 \times 10^8) \text{ joules} \quad (30)$$

(8) We use the Rayleigh Criterion for the angular resolution, $\Delta\theta$,

$$\Delta\theta \text{ (radians)} = \frac{1.22 \lambda}{D} \quad \text{--- (1)} \quad (20)$$

where λ : wavelength

D : diameter of optical telescope

Also, we have

$$s = l (\Delta\theta) \quad \text{--- (2)} \quad (20)$$

where s : diameter of moon crater

l : distance from Moon to Earth

$\Delta\theta$: angular resolution of telescope

From Equation (2), we have

$$\Delta\theta = \frac{2 \times 10^3 \text{ m}}{384,400 \times 10^3 \text{ m}} = 5.203 \times 10^{-6} \text{ radian}$$

Putting this answer into Equation (1),

$$5.203 \times 10^{-6} \text{ radian} = \frac{1.22 \times 555 \times 10^{-9} \text{ m}}{D}$$

(555 nm is the peak of eye sensitivity in the visible light region)

$$\therefore D = \frac{1.22 \times 555 \times 10^{-9} \text{ m}}{5.203 \times 10^{-6} \text{ radian}}$$

$$= 0.130 \text{ m}$$

$$= 5.12 \text{ inches}$$

To resolve the moon crater, we need a telescope of at least 0.130 m diameter

b) We use the Hubble's Law

$$v = H_0 r$$

(40)

where v : recession velocity of galaxy as measured by redshift

H_0 : Hubble constant

r : distance of galaxy to Sun.

Rearranging Hubble's law,

$$\frac{1}{H_0} = \frac{r}{v} = \frac{\text{distance}}{\text{velocity}} = \text{time}$$

∴ by calculating the ~~reciprocal~~ reciprocal of the Hubble constant we can estimate the Age of the Universe.

We have

$$\begin{aligned} \frac{1}{H_0} &= \frac{1}{73.52 \text{ km-s}^{-1} (\text{Mpc})^{-1}} \\ &= \left(\frac{1}{\frac{73.52 \times 10^3 \text{ m-s}^{-1}}{10^6 \times 3.26 \times 3 \times 10^8 \times 365.25 \times 24 \times 3600}} \right) \\ &= \frac{1}{(2.3824 \times 10^{-18} \text{ s}^{-1})} \quad (60) \\ &= 4.1974 \times 10^{17} \text{ s} \\ &= \frac{4.1974 \times 10^{17} \text{ s}}{3600 \times 24 \times 365.25} = 1.330 \times 10^{10} \text{ years} \\ &= 13.30 \text{ billion years} \end{aligned}$$